Parton picture for a strongly-coupled plasma

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Collaboration with Yoshitaka Hatta and Al Mueller (arXiv:0710.2148 and 0710.5297 [hep-th])

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Parton picture for a strongly-coupled plasma - p. 1



Hard probes for sQGP

• QGP just above the phase transition: $T \simeq 2 \div 5 T_c$

- deconfined
- nearly conformal: $(\mathcal{E} 3P)/\mathcal{E} \lesssim 20\%$
- strongly coupled: $\lambda \equiv g^2 N_c \approx 3 \div 6$
 - (RHIC data: v_2 , low viscosity, early thermalization)
- a cousin of the strongly-coupled $\mathcal{N} = 4$ SYM plasma

Motivation

- DIS in pQCD
- DIS off a plasma
- DIS off a BH
- Wave equation
- Low energy
- Saturation momentum
- High energy
- Partons
- Branching



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- AdS/CFT based studies of medium properties ...
 - thermodynamics, hydrodynamics, transport coefficients
- \blacksquare ... and also of 'hard probes' : $\omega,\,q\,\gg\,T$
 - energy loss, momentum broadening, meson dissociation
- How does a strongly-coupled plasma look, when probed on short distances and at high energies ?

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- DIS: the best suited (Gedanken) experiment to measure this

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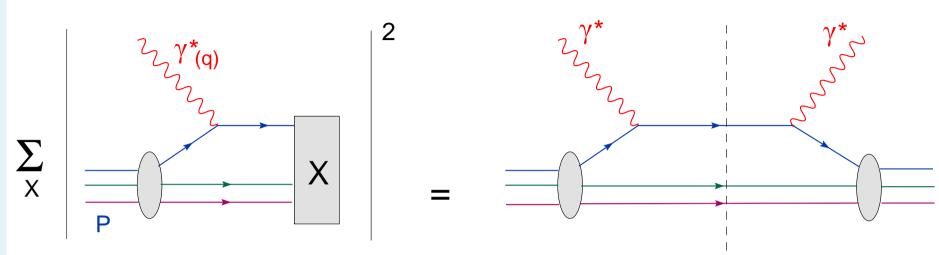
The proton structure functions



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Backup



$$F_{1,2}(x,Q^2) \sim \operatorname{Im} \int \mathrm{d}^4 x \,\mathrm{e}^{-iq \cdot x} \left\langle P \right| \left[J_{\mu}(x), J_{\nu}(0) \right] \left| P \right\rangle$$

Virtual photon absorbed by a quark excitation with

- transverse size $\Delta x_{\perp} \sim 1/Q$, where $Q^2 \equiv -q^{\mu}q_{\mu} \geq 0$
- longitudinal momentum $k_z = xP$, where $x \equiv \frac{Q^2}{2P \cdot q} \simeq \frac{Q^2}{s}$

• $F_2(x, Q^2)$: the quark distribution in the proton

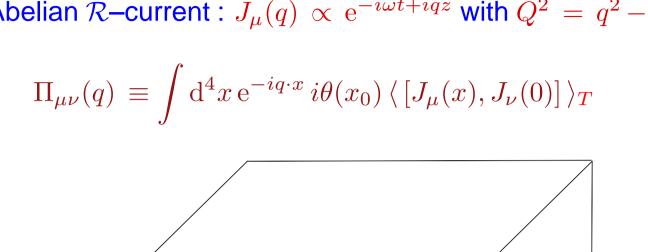


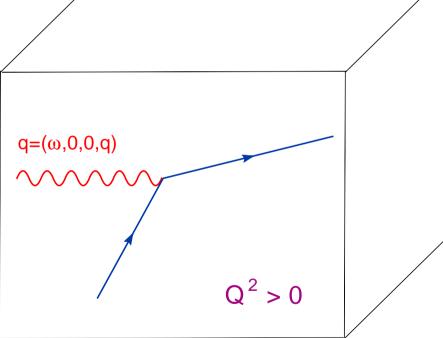
DIS off a $\mathcal{N} = 4$ **SYM plasma**

• An Abelian \mathcal{R} -current : $J_{\mu}(q) \propto e^{-i\omega t + iqz}$ with $Q^2 = q^2 - \omega^2$

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•
$$Q^2 > 0 \Longrightarrow \mathsf{DIS}: F_{1,2}(x,Q^2) \sim \operatorname{Im} \Pi_{\mu\nu}(q)$$



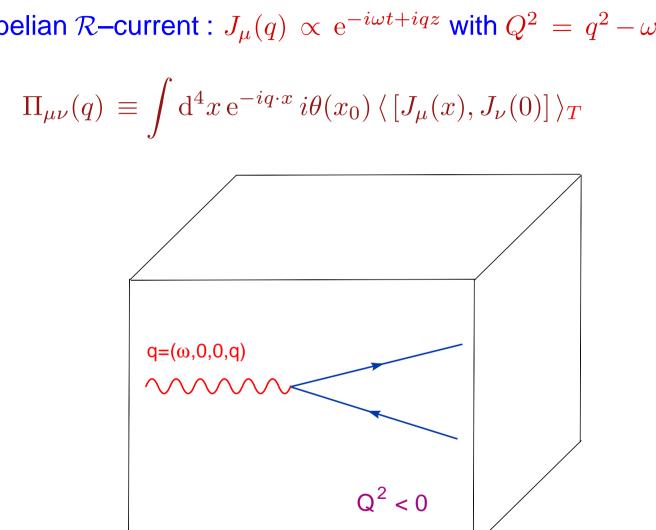
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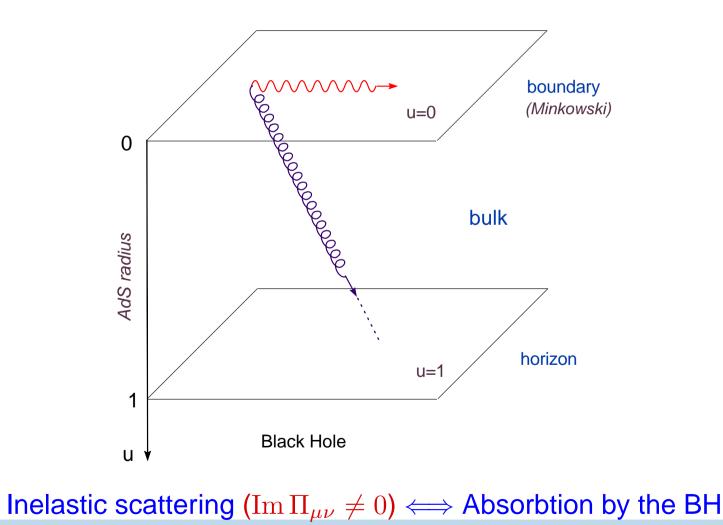
• $Q^2 < 0 \implies$ Jet physics : Energy loss, jet quenching ...



The string dual: AdS_5 **Black Hole**

Strong coupling limit $\lambda \to \infty$ ($N_c \to \infty$) : supergravity

• Metric perturbation in AdS_5 : $A_{\mu}(t, \boldsymbol{x}, u) = e^{-i\omega t + iqz}A_{\mu}(u)$



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● DIS off a BH

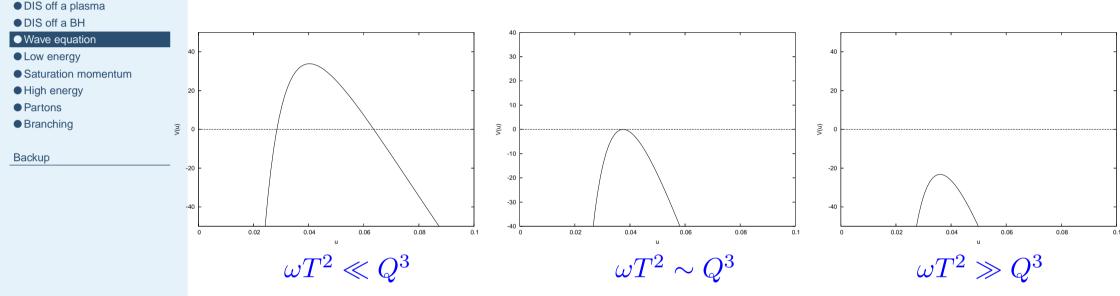
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Motivation
 DIS in pQCD

The gravitational wave equation

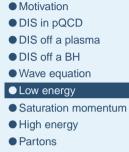
- Einstein equations linearized around AdS₅
- Effective Schroedinger equation: $\psi'' V\psi = 0$



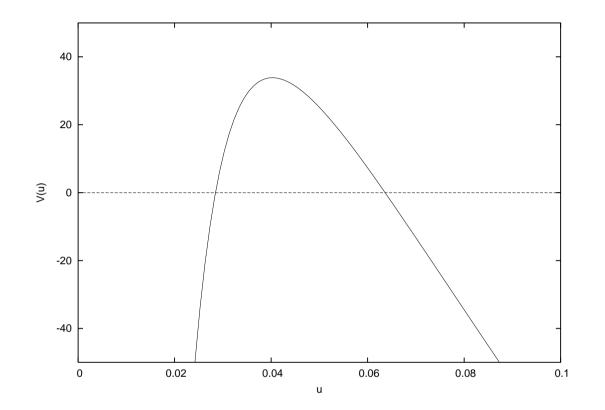
- Repulsive barrier due to the virtuality Q^2
- Attractive potential due to the black hole
- Gravitational interactions are proportional to the energy
- Change of regime when increasing ω and/or T at fixed Q^2



Low energy (or low temperature) : $\omega T^2 \ll Q^3$

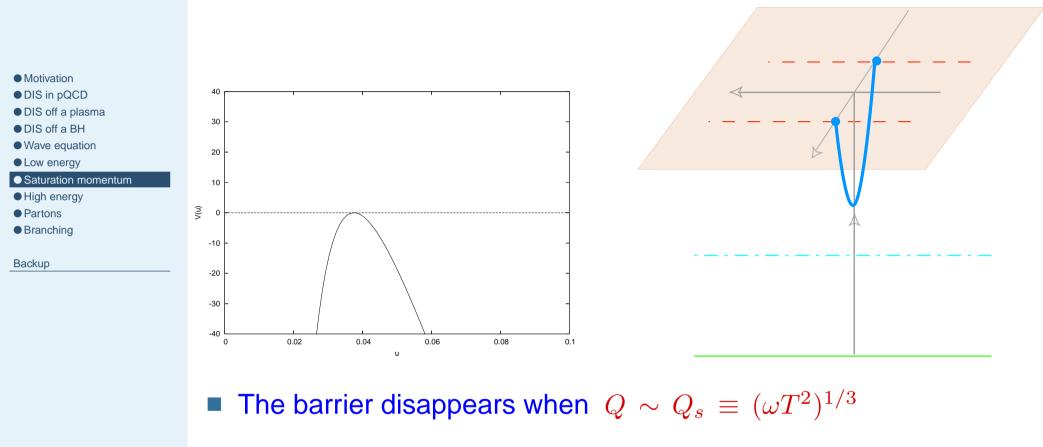


Branching



- Potential barrier: energy-momentum conservation
- Vacuum–like dynamics: a space–like current cannot decay
- Very small imaginary part due to double tunnel effect

Saturation momentum : $\omega T^2 \sim Q^3$



The same as the meson screening length

$$L \sim \frac{1}{Q} \quad \& \quad \gamma \sim \frac{\omega}{Q} \implies L_s \sim \frac{1}{\sqrt{\gamma T}} = \frac{(1-v^2)^{1/4}}{T}$$

[Liu, Rajagopal, Wiedemann; Chernicoff, Garcia, Guijosa; 2006]

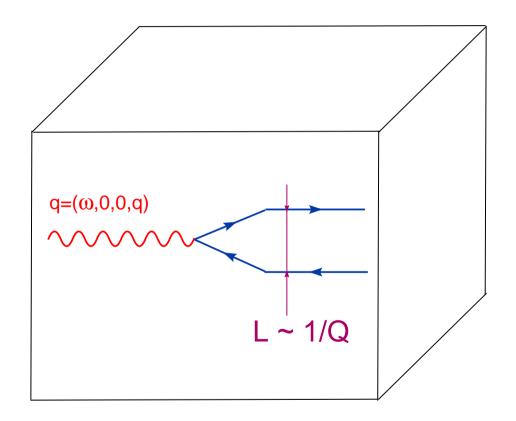


From current to meson

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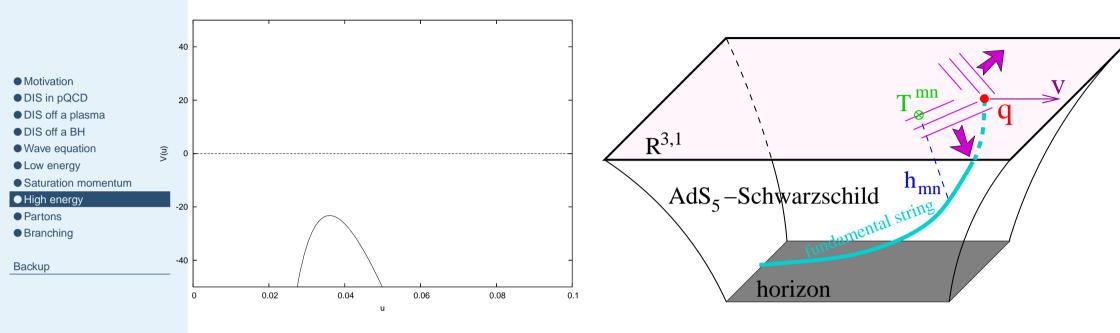
Backup

At very high energy ($\omega \sim q \gg Q$) the current is like a bunch of light–like partons with transverse size $L \sim 1/Q$



The current dynamics can be studied also after the 'meson' breaking a bunch of partons falling in the black hole

High energy (or high temperature) : $\omega T^2 \gg Q^3$



- The wave falls into the $BH \implies$ large imaginary part
- No reflected wave => Total absorption (unitarity limit)

 The line of stationary phase : same as the 'trailing string solution' for a heavy quark (here, with v ~ 1) [Herzog, Yaffe, et al; Gubser et al, 2006]



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Partonic interpretation

• 'Partonic' variables: Q^2 and $x \equiv \frac{Q^2}{2\omega T}$

$$Q_s \simeq (\omega T^2)^{1/3} \iff x_s \simeq \frac{T}{Q}$$

• Low energy $(x \gg x_s)$: $F_2 \simeq 0 \implies$ No large-x partons

High energy ($x \lesssim x_s$): $F_2(x,Q^2) \sim x N_c^2 Q^2$

nterpretation :
$$\frac{1}{x}F_2(x,Q^2) \sim \int^{Q^2} \mathrm{d}^2 k_\perp \frac{\mathrm{d}n}{\mathrm{d}^2 x_\perp \mathrm{d}^2 k_\perp}$$

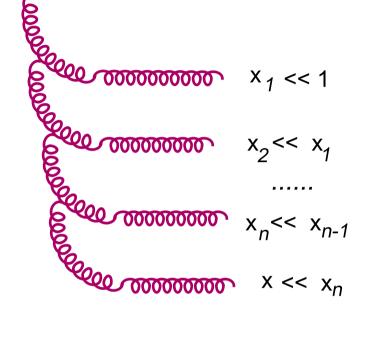
 $\implies \frac{1}{N_c^2} \frac{\mathrm{d}n}{\mathrm{d}^2 x_\perp \mathrm{d}^2 k_\perp} \sim \mathcal{O}(1)$

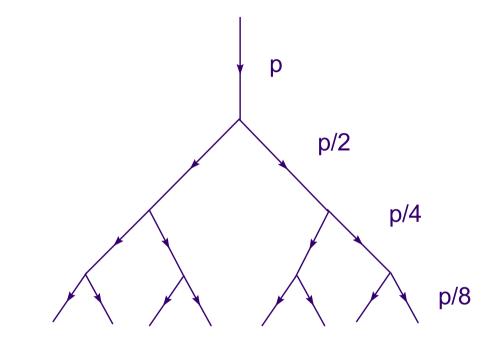
$$\implies$$
 one parton of each color per unit cell in phase–space
 \implies parton saturation (occupation numbers of $\mathcal{O}(1)$)



Parton branching: weak vs. strong coupling

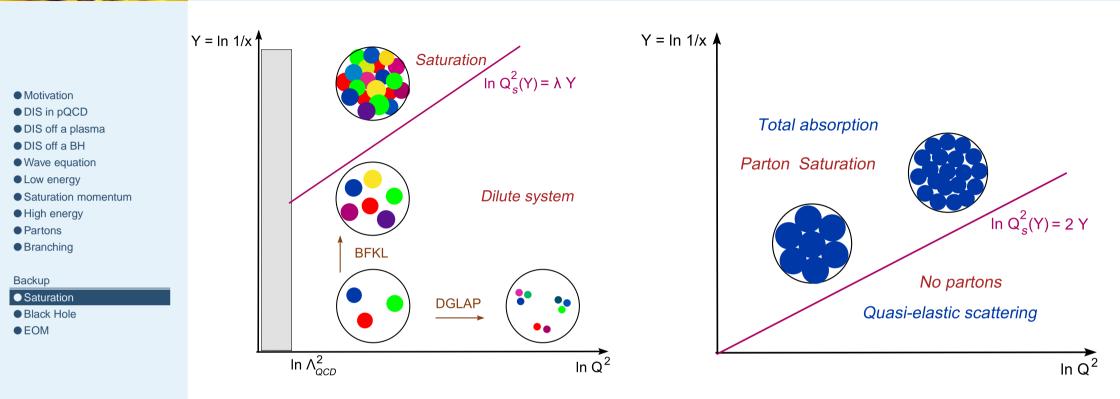
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- Weak coupling: very little energy loss per branching
 - $\bullet\,$ energy is carried by the few remaining partons at large x
- Strong coupling: energy is democratically divided
 - all partons fall down at very small $x \leq x_s \equiv T/Q$
 - the energy is carried by the partons with $x \simeq x_s$





- Saturation exponent : $Q_s^2(x) \propto 1/x^{\lambda} \equiv e^{\lambda Y}$
 - weak coupling (pQCD): $\lambda \approx 1.23 g^2 N_c$ (BFKL Pomeron)
 - strong coupling (plasma): $\lambda = 2$ (graviton)



The string dual of the $\mathcal{N}=4$ SYM plasma

• The $AdS_5 \times S^5$ black hole :

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- Black HoleEOM

 $ds^{2} = \frac{r^{2}}{R^{2}}(-f(r)dt^{2} + dx^{2}) + \frac{R^{2}}{r^{2}f(r)}dr^{2} + R^{2}d\Omega_{5}^{2}$

where $f(r) = 1 - \frac{r_0^4}{r^4}$ and the horizon $r_0 = \pi R^2 T$

- Strong coupling limit $\lambda \to \infty$ ($N_c \to \infty$) : supergravity
- The black hole entropy: $S_{BH} = A/4G$, with A = horizon area

$$\implies \mathcal{S} \equiv \frac{S_{\rm BH}}{V_{3D}} = \frac{\pi^2}{2} N_c^2 T^3 = \frac{3}{4} \mathcal{S}_0$$

• For generic values of $\lambda = g^2 N_c$: $S = f(\lambda) N_c^2 T^3$ where *f* interpolates between 1 and 3/4 \implies Same number of d.o.f. at either weak or strong coupling!



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BackupSaturationBlack Hole

● EOM

Saturation momentum

The wave equations

• Classical EOM (Maxwell eqs. in the AdS_5 BH metric) :

$$\partial_m \left(\sqrt{-g} g^{mn} F_{np} \right) = 0, \qquad F_{mn} = \partial_m A_n - \partial_n A_m$$

Boundary conditions at $r \to \infty$

$$A_{\mu}(t, \boldsymbol{x}, r \to \infty) = e^{-i\omega t + iqz} A_{\mu}, \qquad A_{r}(t, \boldsymbol{x}, r \to \infty) = 0$$

Boundary conditions at the BH horizon r = r₀
 No wave returning from the horizon ⇒ purely outgoing wave
 N.B. The origin of the imaginary part !

The classical action => the current-current correlator:

$$R_{\mu\nu}(q) = \frac{\partial^2 S_{cl}}{\partial A_{\mu} \partial A_{\nu}}$$

• The classical action involves just $(A_{\mu}\partial_{r}A_{\mu})|_{r\to\infty}$