

# <u>Measurement of the open charm</u> <u>cross-section</u>

### in 200 GeV Cu+Cu collisions using STAR @ RHIC

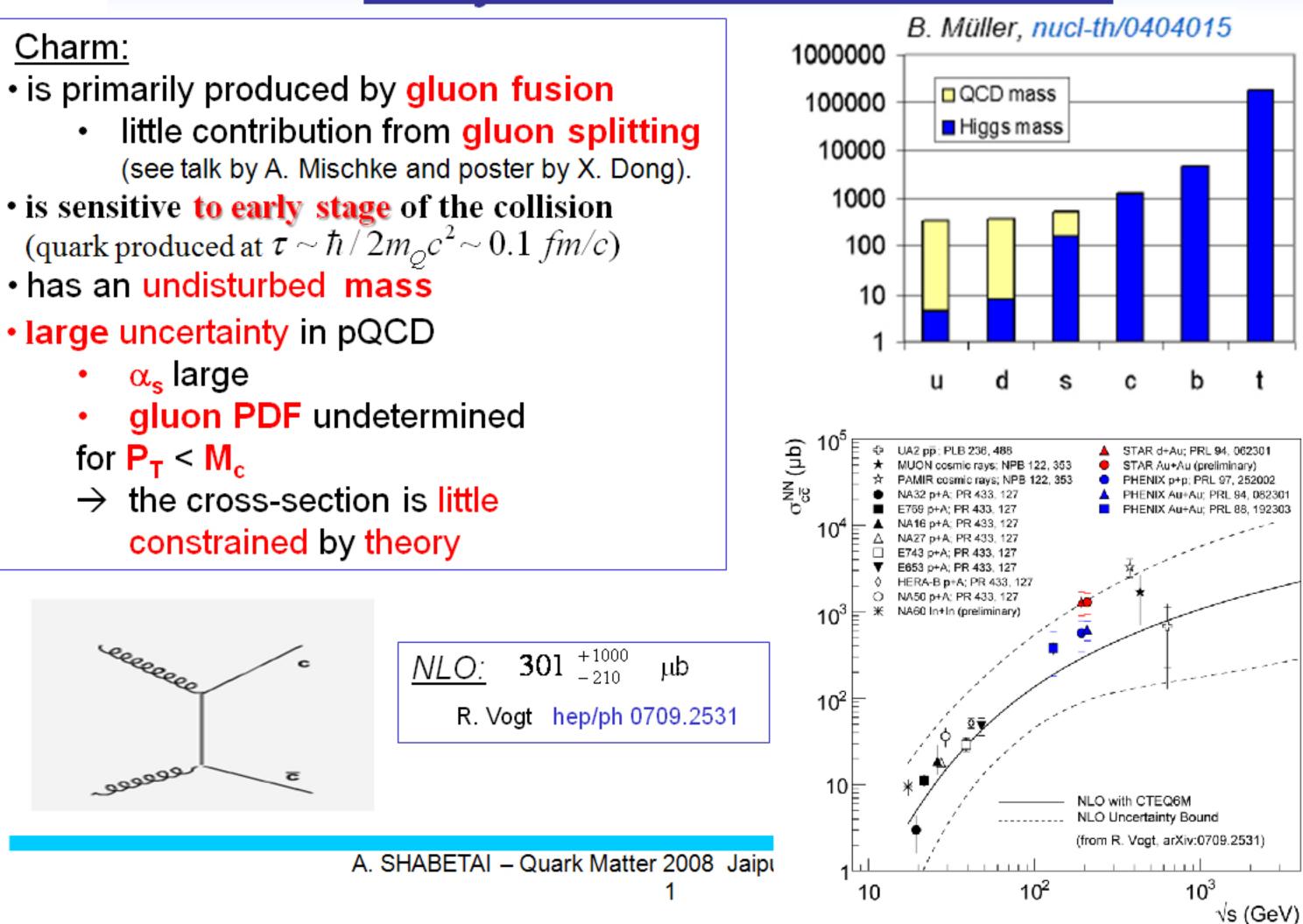
### Alexandre SHABETAI

for the STAR collaboration





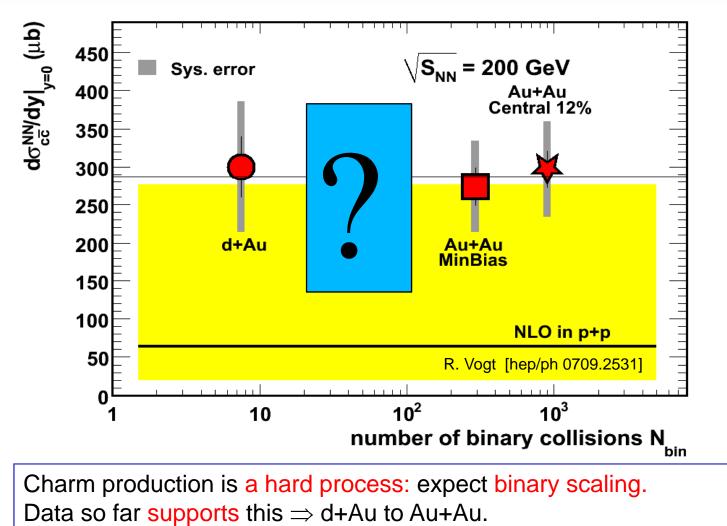
# Physics motivations



**nucl-th/0404015** 



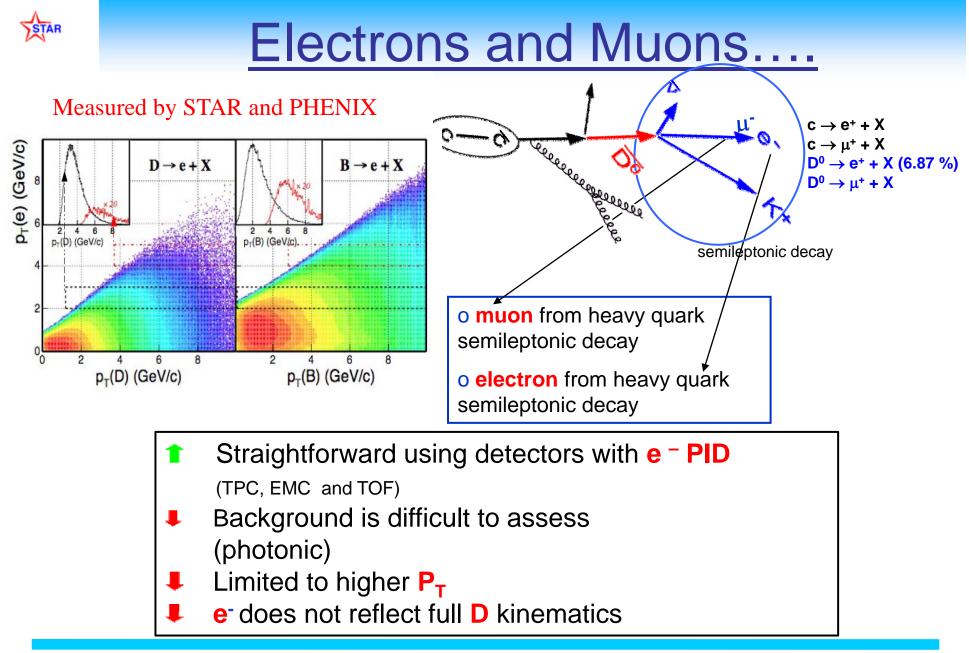
## **Binary scaling**



What about Cu+Cu?

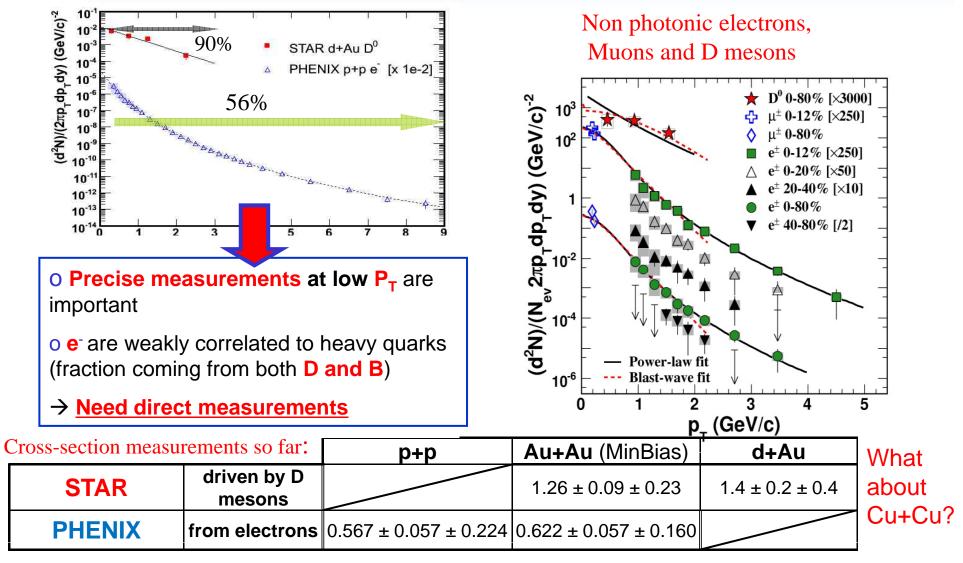


### Open charm measurements via semileptonic decays



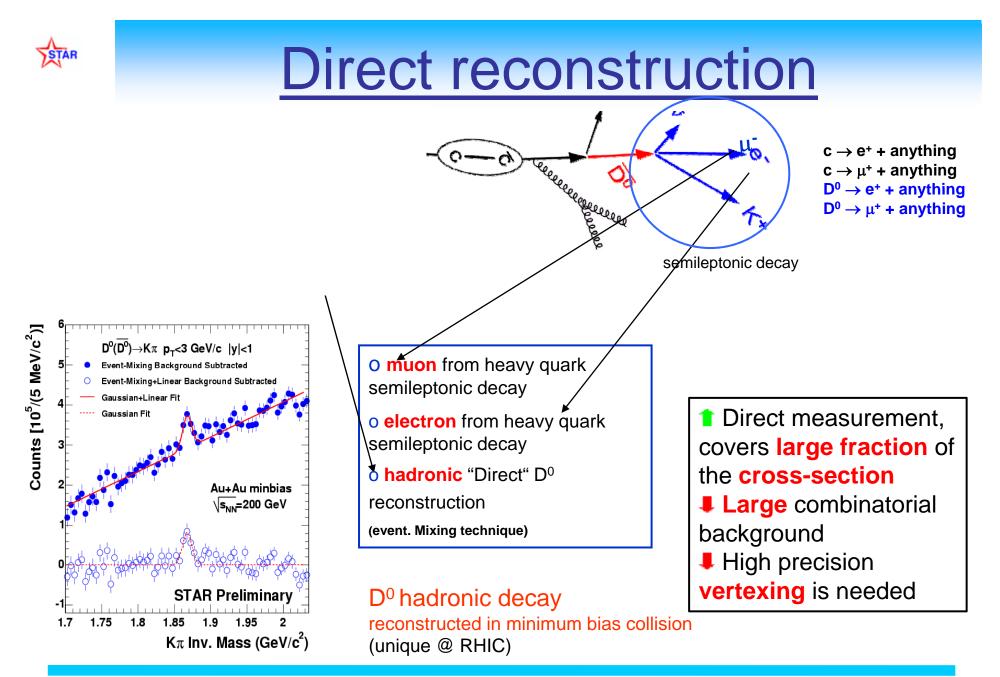


### **Charm Cross-Section to Date**





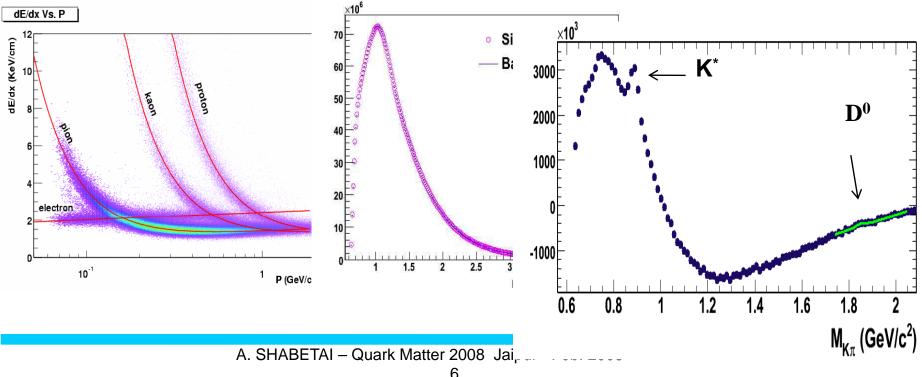
### Open charm measurements via hadronic decays



### STAR

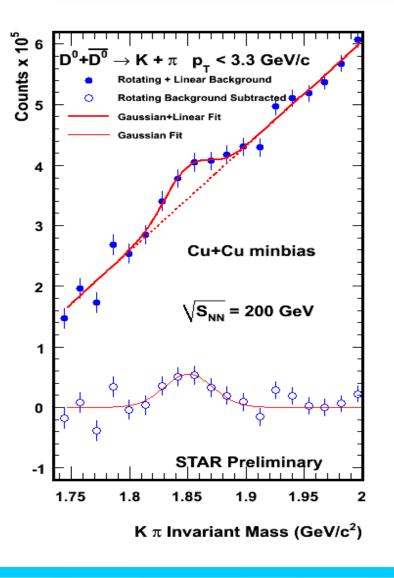
### Direct D<sup>0</sup> reconstruction in STAR

- Pions and Kaons are selected using the TPC
- **\diamond** Combine "same event" pairs  $\Rightarrow$  signal+background
- Combine pairs coming from different events  $\Rightarrow$  background ("mixed events" or "track rotating ")
- ♦ After subtraction  $\Rightarrow$  signal





### Cu+Cu collisions @ 200 GeV



~ 28 Million events used All the statistics available Cu+Cu « minimum bias » (RHIC run V)

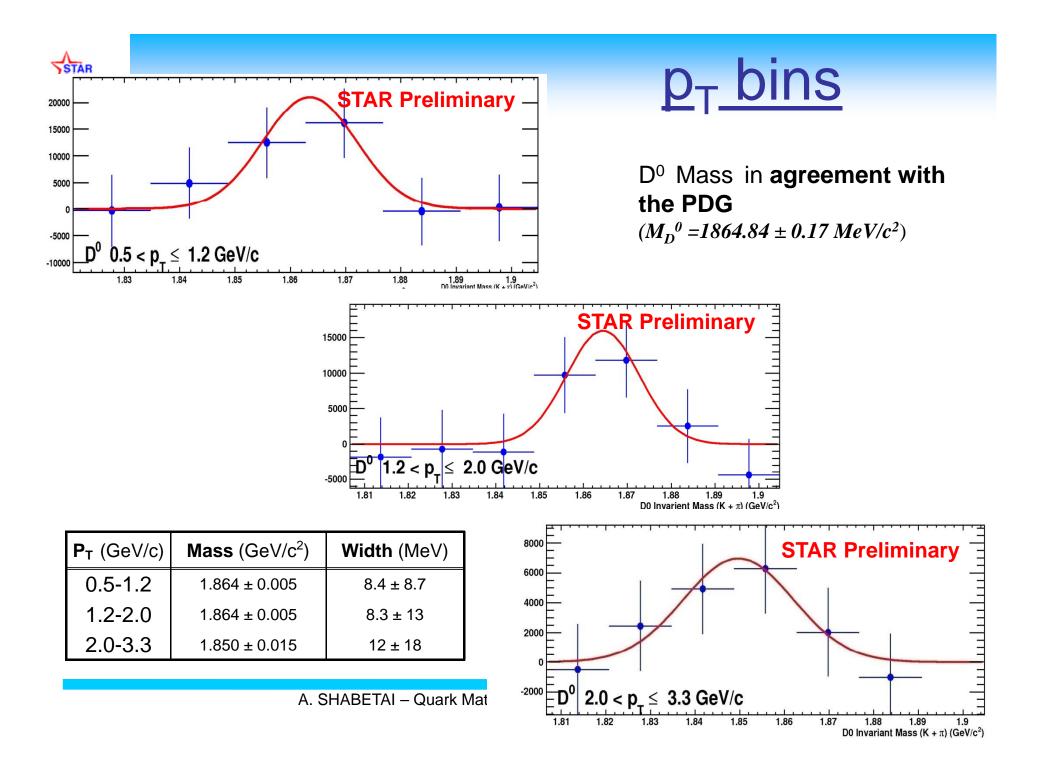
After track rotating or mixed event subtraction: residual background

Low S/B ratio:

$$\frac{S}{B} \approx \frac{1}{600}$$
$$\frac{S}{\sqrt{S+B}} \approx 4$$

Measurement only possible
 because of large S (~ 150 000)
 Large STAR acceptance !

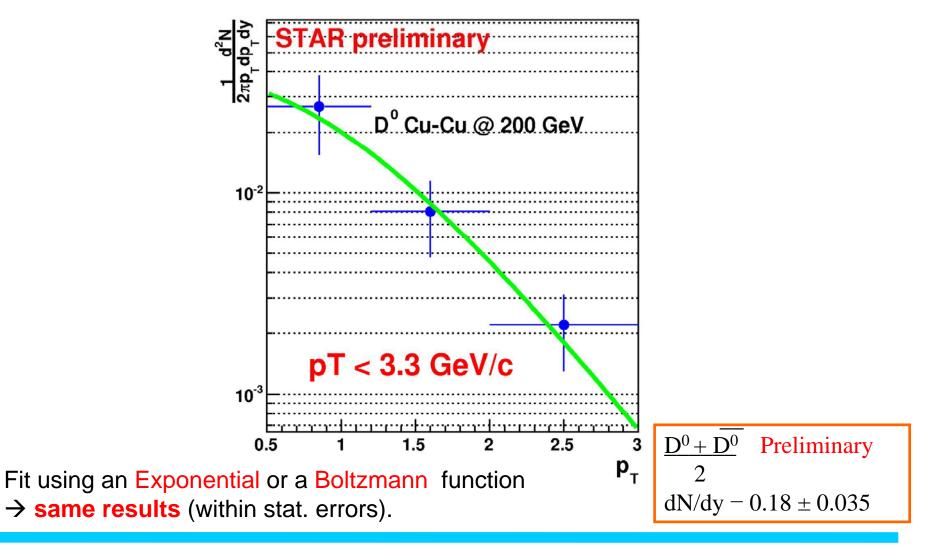
→ Challenging measurement



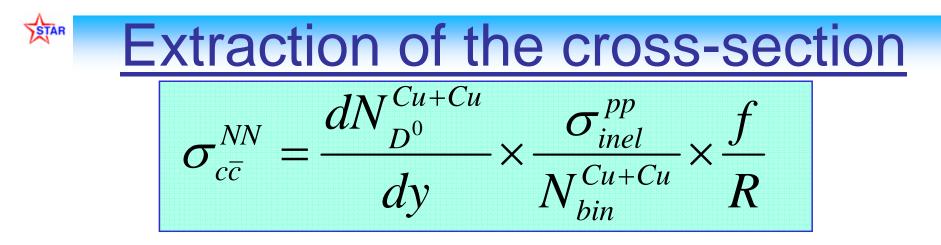




#### **After corrections:**



A. SHABETAI – Quark Matter 2008 Jaipur - Feb. 2008



Number of binary collisions (Glauber)

Inelastic cross-section in p+p (UA5)

Conversion to full rapidity (Pythia)

Ratio obtained from e<sup>+</sup>e<sup>-</sup> collisions

**STAR Preliminary:** 

$$dN_{D^0} / dy = 0.18 + / -0.035 \text{ (stat.)}$$
$$N_{binary}^{Cu+Cu} = 51.5 + 1.0 - 2.9$$

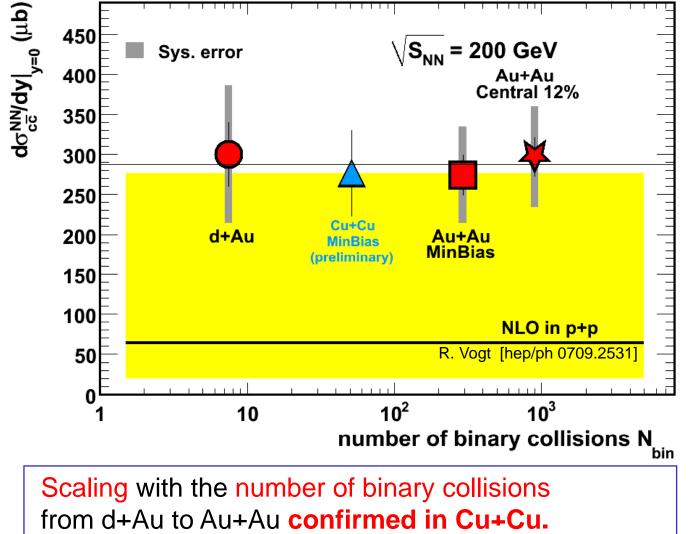
$$\sigma_{inel}^{pp} = 41.8 + / - 0.6 \text{ mb}$$

$$f = 4.7 \pm 0.7$$

$$R = N_{D^0} / N_{c\bar{c}} = 0.54 \pm 0.05$$
$$\Rightarrow \sigma_{c\bar{c}}^{NN} = 1.30 \pm 0.25 \text{ (stat.) mb}$$



# <u>d σ/dy in STAR...</u>



Accurate background subtraction is crucial

Systematic study is ongoing



# **Summary**

#### Today:

- The charm cross-section was measured in Cu+Cu @ 200 GeV ;  $\sigma_{c\bar{c}}^{NN} = 1.30 \pm 0.25 \, mb$
- A direct measurement in Cu-Cu is consistent with a scaling of the crosssection with N<sub>bin</sub> (at low pT).
- Theory: large uncertainty in pQCD calculations and data points are needed.

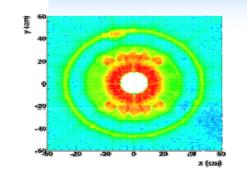
#### In the Future:

- STAR low material runs
- use of SSD/SVT and eventually the HFT upgrade (2010-2012)

will allow:

- ✤ precise measurements of the charm cross-section
- direct topological measurements of charm and of its anisotropy parameter
   V<sub>2</sub>, R<sub>AA</sub>, R<sub>CP</sub>
- isolation of the bottom contributions





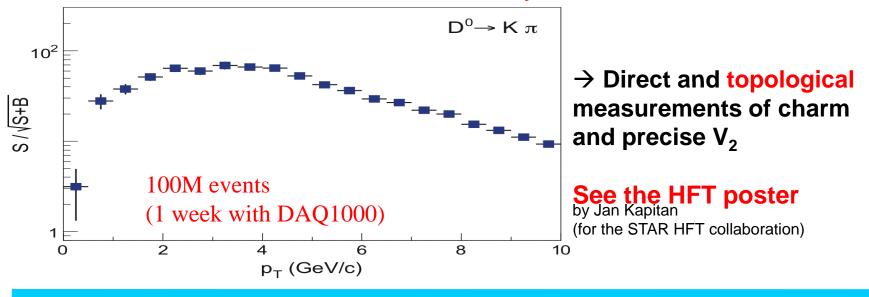
### <u>Outlook</u>

- Low material run (without the SVT/SSD)
- $\rightarrow$  low radiation length in run VIII
- ightarrow reduce the photonic background

Reconstructing the secondary vertex with SVT/SSD in Au+Au (run VII)

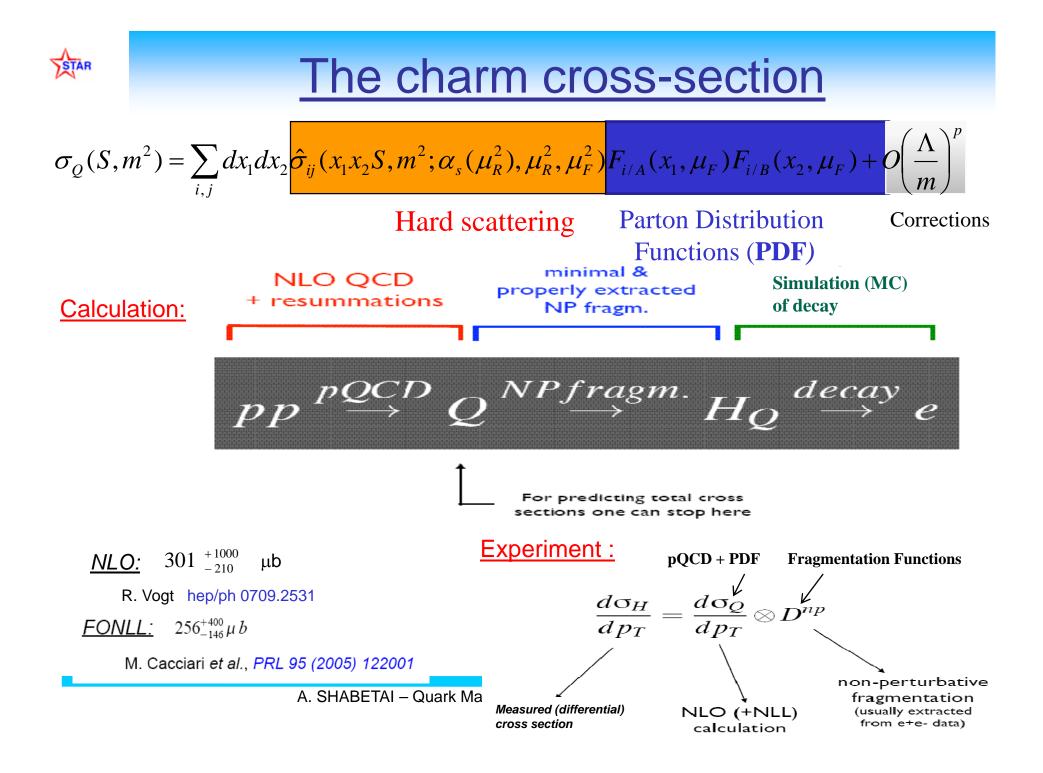
#### ✤ « Upgrade » for RHIC2

and especially The future STAR *Heavy Flavor Tracker* 











# Cross-section – How well is the calculation constrained?

- Energy
- Charm quark mass (m<sub>c</sub>)

### - Scales

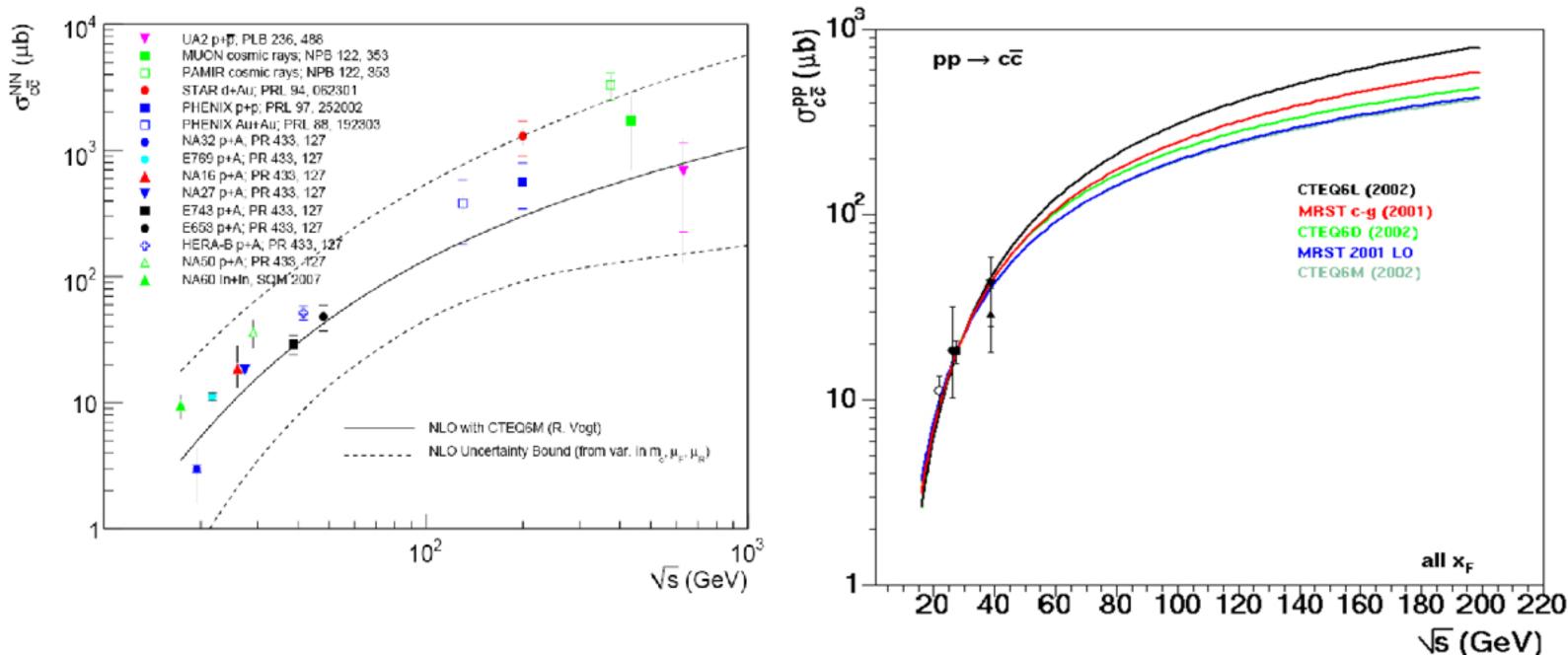
- -m<sub>R</sub>: fragmentation scale
- -m<sub>E</sub> : factorization scale
- -a<sub>s</sub>: strong coupling
- PDF used

### Example:

FONLL:
$$\mu_F = \mu_R = \mu = \sqrt{p_T^2}$$
PYTHIA:CTEQ5M1, MSEL=1NLO:MRST $\mu = 2m_c$ 

- Fragmentation Functions (FF) non perturbative inserted in a







 $+ m_c^2, m_c = 1.2 GeV/c^2$ 

### $, m_c = 1.2 GeV/c^2$

H. Wöhri and C. Lourenço Jphys G Nucl Part Phys 30 (2004)315

## How to compare measurements to calculations ?

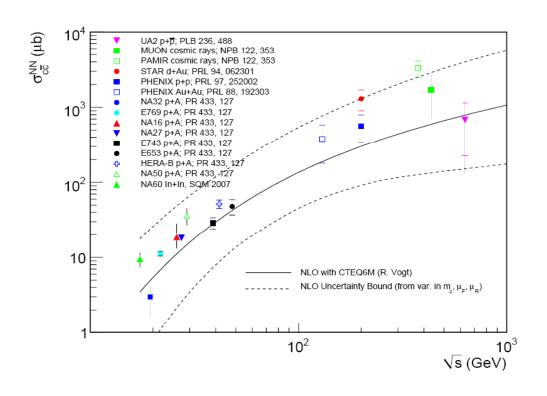
Using QCD (and pQCD) :

STAR

- Heavy flavor cross-section can be correctly predicted

Differential cross-section (as function of momentum, rapidity or energy...), requires « adding a <u>minimal</u>, <u>self-consistent and universal</u> set of non perturbative input parameters »

#### Matteo Cacciari ISMD 2007



To make an accurate comparison, one should:

- Use **dedicated theoretical tools** (FONLL and now NNLO)

- Use **adequate parameters** (mass, renormalization and factorization scales, coupling), Partons Distribution Functions (PDF) and Fragmentation Functions (FF).

- <u>Minimize extrapolations and</u> <u>deconvolutions between</u> <u>measurements and theory</u>

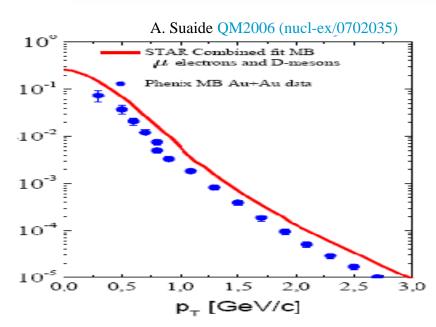
→ If and only if all those conditions are satisfied, a good agreement between measurements and calculations can be reached

 $\rightarrow$ In real life, the error band is large and data points are needed...

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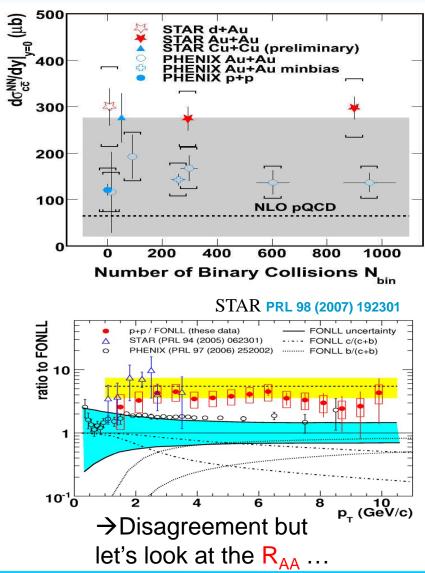


### PHENIX vs. STAR



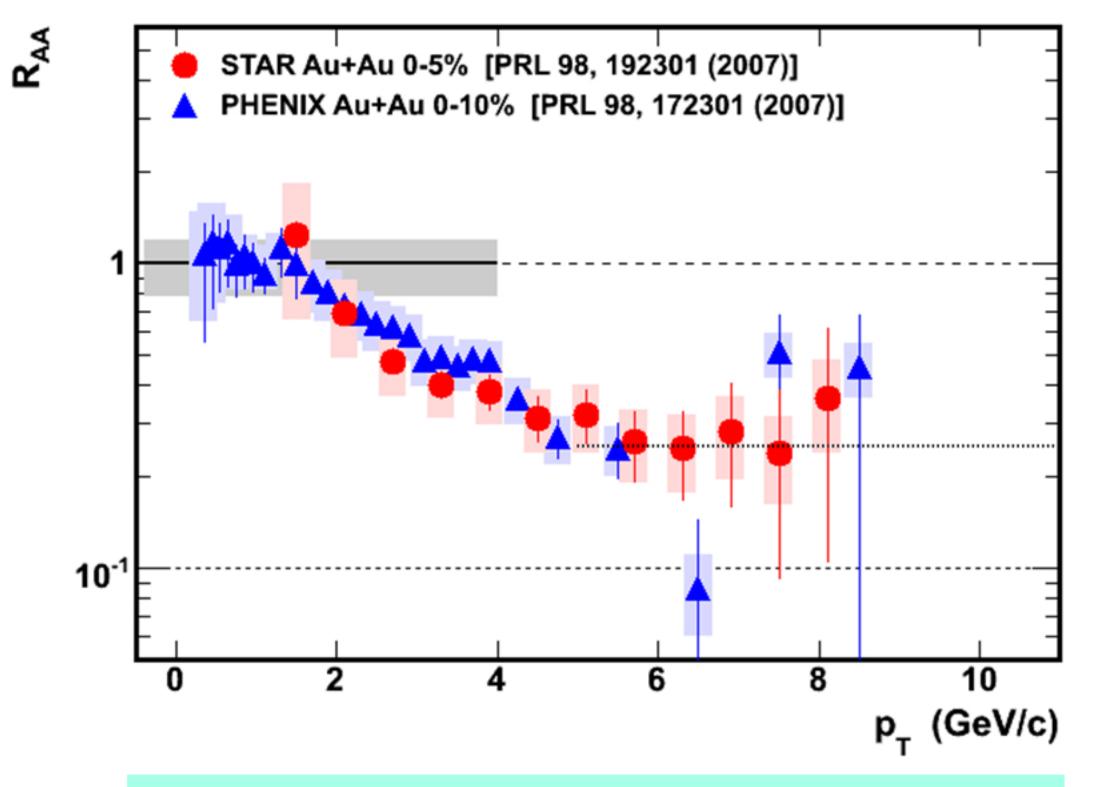
- Spectra shapes are the same. STAR and PHENIX are seeing the same scaling with Nbin.

The value of the cross-section
is not the same (factor 2 to 3)
STAR and PHENIX are both above FONLL predictions...





# R<sub>AA</sub>: (e<sup>-</sup>) from d+Au to Au+Au central



 $R_{AA}$  in agreement between STAR and PHENIX  $\rightarrow$  Is there a normalization issue?

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Dead cone effect not observed ... (non photonic e<sup>-</sup>)